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# Nonlocally-Correlated Disorder and Delocalization in One Dimension: Density of States

Ikuo Ichinose<sup>†1</sup> and Masaomi Kimura<sup>\*2</sup>

<sup>†</sup>Institute of Physics, University of Tokyo, Komaba, Tokyo, 153-8902 Japan

<sup>\*</sup>Institute for Cosmic Ray Research, University of Tokyo, Tanashi, Tokyo, 188-8502  
Japan

## Abstract

We study delocalization transition in a one-dimensional electron system with quenched disorder by using supersymmetric (SUSY) methods. Especially we focus on effects of nonlocal correlation of disorder, for most of studies given so far considered  $\delta$ -function type white noise disorder. We obtain wave function of the “lowest-energy” state which dominates partition function in the limit of large system size. Density of states is calculated in the scaling region. The result shows that delocalization transition is stable against nonlocal short-ranged correlation of disorder. Especially states near the band center are enhanced by the correlation of disorder which partially suppresses random fluctuation of disorder. Physical picture of the localization and the delocalization transition is discussed.

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<sup>1</sup>e-mail address: ikuo@hep1.c.u-tokyo.ac.jp

<sup>2</sup>e-mail address: masaomi@ctpc1.icrr.u-tokyo.ac.jp

# 1 Introduction

Quenched disorder plays important roles in various physical phenomena. Anderson localization is one of these examples[1]. Especially in two and lower dimensions almost all states are localized, and extended delocalized states are isolated points in physical parameter region. The transition between quantum Hall plateaus is such an example.

For the transition between integer quantum Hall plateaus, useful model, named network model, has been proposed[2], which incorporates effects of localization and quantum tunneling in a strong magnetic field. Numerical studies indicate that the network model belongs to the same universality class of the transition between integer quantum Hall plateaus. However it is rather difficult to solve the network model analytically because there exists no controllable parameter for perturbative expansion and also because of nature of quenched disorder itself. Compared with the two-dimensional (2D) systems, one dimensional (1D) systems are more tractable. These 1D systems include Dyson's study on random strength harmonic springs[3], random Ising model by McCoy and Wu[4], and more recently random exchange spin chains[5] and random hopping tight-binding (RHTB) model[6, 7, 8].

Recently supersymmetric (SUSY) methods appear useful for handling the quenched disorder. SUSY methods are applied to the network model[9], the 1D RHTB model[7] etc. In this paper we shall revisit the 1D RHTB model by applying the SUSY methods. Especially we shall consider *nonlocally correlated* quenched disorder and study stability of the delocalization transition which exists in the case of the  $\delta$ -function-type white noise disorder. The model contains two parameters which control magnitude of fluctuation and correlation length of disorder. We expect that we can get detailed physical picture of (de)localization transition from calculations of density of states, Green's functions, etc. as a function of the above parameters. Another motivation of the present work is rather technical, i.e., we show how to use the SUSY methods for nonlocally correlated disorder systems.

Hamiltonian of the RHTB model is given by

$$\mathcal{H}_{RHTB} = - \sum_n t_n (c_n^\dagger c_{n+1} + c_{n+1}^\dagger c_n), \quad (1.1)$$

where  $c_n$  is annihilation operator of spinless fermion at site  $n$  and  $t_n$ 's are random hopping parameters. As we shall see, staggered part of fluctuation of  $t_n$ 's plays an important role at low energies and generates energy gap for most of fermion modes. Extended excitations are located near the band center, and they are described by a Dirac fermion with randomly varying mass.

In the SUSY methods, bosonic variables are introduced. In terms of them, density of states, Green's functions, etc are expressed in a compact way. Average over random variables of *nonlocal* correlation can be performed by introducing another bosonic variable, and because of that there appear couplings between the fermion and its bosonic SUSY partner. The model is reduced to a quantum mechanical system of one fermionic and two bosonic variables. The lowest-energy state which dominates the partition function is annihilated by SUSY charges. We obtain wave function of the lowest-energy state and calculate density of states near the band center.

This paper is organized as follows; In Sect.2, we shall review the RHTB model, its continuum limit and SUSY methods which are applied for nonlocally correlated disorder. Most of our notations follow those by Balents and Fisher in Ref.[7]. In Sect.3, transfer Hamiltonian is obtained regarding the spatial coordinate as time. Then the system is reduced to a quantum mechanical system. In Sect.4, we shall obtain an equation of motion of the ground state, which dominates the partition function and density of states. Section 5 is devoted for solution to equation of motion, and in Sect.6 by using the ground state solution we shall calculate density of states of fermions. Section 7 is devoted for discussion.

In this paper, we calculate density of states of fermions. We shall report Green's functions in a future publication[10]. These calculations are quite useful for understanding localization and the delocalization transition.

## 2 Continuum limit and SUSY

The random hopping parameters fluctuate as

$$t_n = t + \delta t_n, \quad (2.1)$$

where  $t$  is some finite constant and we normalize as  $t = \frac{1}{2a_l}$  and  $a_l$  is the lattice spacing. We assume that r.m.s. of random variables  $\delta t_n$  is sufficiently small compared with the band width  $t$ . Continuum limit of the system (1.1) is then easily obtained. We shall focus on modes near the band center because delocalization transition occurs there. Excitations near the band center of fermions are described by smoothly varying right and left-moving fermion fields  $\psi_R$  and  $\psi_L$ ,

$$c_n = e^{ik_F n} \psi_R(x) + e^{-ik_F n} \psi_L(x), \quad k_F = \frac{\pi}{2}, \quad (2.2)$$

where continuous coordinate  $x = na_l$ .

The most relevant part of the random hopping parameters  $\delta t_n$  is their staggered part,  $\delta t_n \sim (-1)^n m(x)$ . Then in terms of  $\psi_R(x)$  and  $\psi_L(x)$ , we obtain continuum limit of the Hamiltonian (1.1) as

$$\begin{aligned} \mathcal{H}_c &= - \int dx \left[ \psi_R^\dagger i \partial_x \psi_R - \psi_L^\dagger i \partial_x \psi_L - im(x)(\psi_R^\dagger \psi_L - \psi_L^\dagger \psi_R) \right], \\ &= - \int dx \psi^\dagger h \psi, \\ h &= -i\sigma^z \partial_x + m(x)\sigma^y, \quad \psi = (\psi_R \psi_L)^t. \end{aligned} \quad (2.3)$$

The random mass  $m(x)$  is decomposed into a uniform and random piece as

$$m(x) = m_0 + \phi(x), \quad (2.4)$$

where  $[\phi]_{ens} = 0$  and

$$[\phi(x)\phi(y)]_{ens} = \frac{g}{2\lambda} \exp(-|x-y|/\lambda), \quad (2.5)$$

with positive parameters  $g$  and  $\lambda$ . It is easily verified

$$\int dx [\phi(x)\phi(y)]_{ens} = g,$$

and

$$[\phi(x)\phi(y)]_{ens} \rightarrow g\delta(x-y) \text{ as } \lambda \rightarrow 0.$$

It is obvious that the parameter  $g$  controls magnitude of fluctuation and  $\lambda$  is the correlation length of disorder.

Single fermion Green's function at energy  $\omega$  is defined as

$$G_{\alpha\beta}(x, y; i\omega) = \langle x, \alpha | \frac{1}{h - i\omega} | y, \beta \rangle, \quad \alpha, \beta = R, L, \quad (2.6)$$

where  $|x, \alpha\rangle$  is a *normalized* position eigenstate of fermion at  $x$  and chirality  $\alpha$ . By functional integral,

$$\begin{aligned} G_{\alpha\beta}(x, y; i\omega) &= i\langle \psi_\alpha(x) \bar{\psi}_\beta(y) \rangle_\psi \\ &= \frac{1}{Z_\psi} \int \mathcal{D}\psi \mathcal{D}\bar{\psi} \psi_\alpha(x) \bar{\psi}_\beta(y) e^{-S_\psi}, \\ S_\psi &= \int dx \bar{\psi}(ih + \omega)\psi, \\ Z_\psi &= \int \mathcal{D}\psi \mathcal{D}\bar{\psi} e^{-S_\psi}. \end{aligned} \quad (2.7)$$

Ensemble averaged Green's function is obtained from (2.6) as

$$\overline{G}_{\alpha\beta}(x - y; i\omega) = \left[ \langle x, \alpha | \frac{1}{h - i\omega} | y, \beta \rangle \right]_{ens}, \quad (2.8)$$

where the ensemble average is taken with respect to  $\phi(x)$  in (2.3) and (2.4) according to (2.5). To take the ensemble average seems formidable. However by introducing SUSY partner, Green's function  $\overline{G}_{\alpha\beta}(x - y; i\omega)$  etc can be expressed by a functional integral in a compact way. Point is that integration over the bosonic SUSY partner cancels the fermionic determinant and normalization of the fermionic states is automatically taken into account. Introducing the bosonic superpartner  $\xi$ ,

$$\overline{G}_{\alpha\beta}(x - y; i\omega) = i\langle \psi_\alpha(x) \bar{\psi}_\beta(y) \rangle_S, \quad (2.9)$$

where

$$\langle \mathcal{A} \rangle_S = \left[ \int \mathcal{D}\psi \mathcal{D}\bar{\psi} \mathcal{D}\xi \mathcal{D}\bar{\xi} \mathcal{A} e^{-S} \right]_{ens}, \quad (2.10)$$

with

$$S = \int dx \left[ \bar{\psi}(ih + \omega)\psi + \bar{\xi}(ih + \omega)\xi \right]. \quad (2.11)$$

The ensemble average in (2.10) can be performed in the following way. We first notice identity such as

$$(-\lambda^2 \partial_x^2 + 1) \frac{1}{2\lambda} \exp(-|x - y|/\lambda) = \delta(x - y). \quad (2.12)$$

Then the ensemble average can be converted into the functional integral form,

$$[\phi(x)\phi(y)\dots]_{ens} = \int \mathcal{D}\phi(x') (\phi(x)\phi(y)\dots) \exp(-S_\phi[\phi(x')]), \quad (2.13)$$

$$S_\phi[\phi(x)] = \int dx \frac{1}{4g} \phi(x) (-\lambda^2 \partial_x^2 + 1) \phi(x). \quad (2.14)$$

From (2.10), (2.11) and (2.13), the expectation value of operator  $\mathcal{A}$  is given by

$$\langle \mathcal{A} \rangle_S = \int \mathcal{D}\psi \mathcal{D}\bar{\psi} \mathcal{D}\xi \mathcal{D}\bar{\xi} \mathcal{D}\phi \mathcal{A} \exp(-(S + S_\phi)). \quad (2.15)$$

The above total action is obviously SUSY under  $\psi \leftrightarrow \xi$  and has such a form that the SUSY partners  $\psi$  and  $\xi$  couple to the “dynamical” real scalar field  $\phi$  which is “singlet” under SUSY transformation.

In the following section, we shall obtain a transfer Hamiltonian by regarding the spatial coordinate  $x$  as time. The system reduces to a quantum mechanical system with the two bosonic and one fermionic variables.

### 3 Transfer Hamiltonian

In this section, we shall obtain a transfer Hamiltonian by regarding the spatial coordinate  $x$  as time in the functional integral representation (2.15). Then the system reduces to a quantum mechanical system. In Ref.[7], the following canonical creation and annihilation operators are introduced corresponding to the functional integral variables,

$$\begin{aligned} \psi_R &\rightarrow F_\uparrow, & \bar{\psi}_R &\rightarrow F_\uparrow^\dagger, & \psi_L &\rightarrow -F_\downarrow, & \bar{\psi}_L &\rightarrow F_\downarrow^\dagger, \\ \xi_R &\rightarrow B_\uparrow, & \bar{\xi}_R &\rightarrow B_\uparrow^\dagger, & \xi_L &\rightarrow -B_\downarrow, & \bar{\xi}_L &\rightarrow -B_\downarrow^\dagger. \end{aligned} \quad (3.1)$$

It is useful to define fermionic and bosonic spin operators,

$$\begin{aligned}\vec{\mathcal{J}}_f &= \frac{1}{2}F^\dagger \vec{\sigma} F, \\ \vec{\mathcal{J}}_b &= \frac{1}{2}\bar{B} \vec{\sigma} B,\end{aligned}\tag{3.2}$$

where

$$\bar{B} = B^\dagger \sigma^z.\tag{3.3}$$

It is proved that  $\vec{\mathcal{J}}_f$  and  $\vec{\mathcal{J}}_b$  satisfy  $SU(2)$  and  $SU(1,1)$  algebras, respectively, and they commute with SUSY charges  $Q$  and  $\bar{Q}$ ,

$$Q = \bar{B}F, \quad \bar{Q} = F^\dagger B.\tag{3.4}$$

Transfer Hamiltonian of  $F_\sigma$  and  $B_\sigma$  ( $\sigma = \uparrow, \downarrow$ ) part is written in terms of the spin operators.

Transfer Hamiltonian of  $\phi$  is also obtained from (2.14). The system of  $\phi$  is nothing but a simple harmonic oscillator linearly coupled with the SUSY spin  $\mathcal{J} = \mathcal{J}_f + \mathcal{J}_b$ . In terms of the spin operators and the canonical boson operators of the harmonic oscillator  $a$ ,  $a^\dagger$  which correspond to  $\phi$ , Hamiltonian of the system is given as,

$$H = 2\omega \mathcal{J}^z + 2m_0 \mathcal{J}^x + \sqrt{\frac{4g}{\lambda}} \mathcal{J}^x (a + a^\dagger) + \frac{1}{\lambda} (a^\dagger a + \frac{1}{2}).\tag{3.5}$$

Fermionic states of  $F_\sigma$  are specified by representations of  $SU(2)$ . Similarly bosonic states of  $B_\sigma$  form multiplet of irreducible representations of  $SU(1,1)$ , which are specified by total spin

$$J^2 = (N_B^2 + 2N_B)/4, \quad N_B = \bar{B}B = B_\uparrow^\dagger B_\uparrow - B_\downarrow^\dagger B_\downarrow,\tag{3.6}$$

and  $z$ -component of spin  $J^z$ , i.e.,

$$\begin{aligned}J^2 |jn\rangle &= j(j+1) |jn\rangle, \\ J^z |jn\rangle &= \left[ \frac{1 + |2j+1|}{2} + n \right] |jn\rangle,\end{aligned}\tag{3.7}$$

where it should be remarked here that the total spin takes half integers, namely,  $j = 0, \pm 1/2, \pm 1, \dots$ .

In Ref.[7], structure of the above quantum mechanical states of the SUSY partners is studied in detail and it is shown that SUSY-invariant states are explicitly given as follows;

$$|n\rangle_0 = \begin{cases} \frac{1}{\sqrt{2}}[|-1/2, n\rangle \otimes |\downarrow\rangle_F + |-1/2, n-1\rangle \otimes |\uparrow\rangle_F], & n > 0 \\ |-1/2, 0\rangle \otimes |\downarrow\rangle_F, & n = 0 \end{cases} \quad (3.8)$$

where  $|\sigma\rangle_F$  is the fermionic state of spin  $\sigma$ .

## 4 The ground state

From the functional-integral representation (2.15), it is obvious that the partition function  $Z = \text{Tr}((-1)^{N_f} e^{-LH})$ , where  $N_f$  = fermion number and  $L$  = system size, acquires contribution only from the scalar field  $\phi$  with the action  $S_\phi$  because fermionic and bosonic determinants cancel with each other for the SUSY partners. Then in the large system size limit  $L \rightarrow \infty$ ,  $Z = e^{-L/2\lambda}$ . The “ground state” which satisfies

$$H|0\rangle = \frac{1}{2\lambda}|0\rangle, \quad (4.1)$$

or

$$\left(2\omega\mathcal{J}^z + 2m_0\mathcal{J}^x + \sqrt{\frac{2}{\lambda}}\mathcal{J}^x(a + a^\dagger) + \frac{1}{\lambda}a^\dagger a\right)|0\rangle = 0, \quad (4.2)$$

dominates the partition function and the density of states of fermions which we shall calculate in this paper. In Eq.(4.2), we have rescaled the parameters as  $\omega \rightarrow 2g\omega$ ,  $m_0 \rightarrow 2gm_0$  and  $\lambda \rightarrow \frac{\lambda}{2g}$ . This rescaling leads the general Hamiltonian to the one of  $g = \frac{1}{2}$ . After calculation, we shall scale back the above parameters and recover dependence on  $g$ .

The SUSY vacuum should be invariant under SUSY transformation and it must be annihilated by both  $Q$  and  $\bar{Q}$ , and therefore it is given by a sum of a direct product



of  $|n\rangle_0$  in Eq.(3.8) and  $|m\rangle_H$ , which is the eigenstates of the harmonic oscillator, i.e. ;

$$|0\rangle = \sum_{n,m} \phi_{n,m} |n\rangle_0 |m\rangle_H, \quad (4.3)$$

where  $|m\rangle_H = \frac{1}{\sqrt{m!}}(a^\dagger)^m |0\rangle_H$  with the “vacuum” of the harmonic oscillator  $|0\rangle_H$ . Norm of the state (4.3) is given by<sup>1</sup>

$$\langle 0|0\rangle = \sum_m |\phi_{0,m}|^2. \quad (4.4)$$

Let us insert (4.3) into (4.2),

$$\begin{aligned} & \sum_{n,m} \phi_{n,m} \left( 2\omega n |n\rangle_0 |m\rangle_H + m_0 ((n+1)|n+1\rangle_0 - (n-1)|n-1\rangle_0) |m\rangle_H \right. \\ & + \sqrt{\frac{1}{2\lambda}} ((n+1)|n+1\rangle_0 - (n-1)|n-1\rangle_0) (\sqrt{m+1}|m+1\rangle_H + \sqrt{m}|m-1\rangle_H) \\ & \left. + \frac{1}{\lambda} m |n\rangle_0 |m\rangle_H \right) = 0. \end{aligned} \quad (4.5)$$

Rearranging the indices  $n$  and  $m$  of the subscripts of  $\phi_{n,m}$  and the labels of states, we obtain equation of  $\phi_{n,m}$  for the vacuum,

$$\begin{aligned} & (2\omega n + \frac{1}{\lambda} m) \phi_{n,m} + m_0 \cdot n \cdot (\phi_{n-1,m} - \phi_{n+1,m}) \\ & + \sqrt{\frac{1}{2\lambda}} n (\sqrt{m+1} (\phi_{n-1,m+1} - \phi_{n+1,m+1}) + \sqrt{m} (\phi_{n-1,m-1} - \phi_{n+1,m-1})) = 0. \end{aligned} \quad (4.6)$$

Equation (4.6) for  $n = 0$  can be solved easily and the solution is

$$\phi_{0,m} = \begin{cases} 0, & m \neq 0, \\ 1, & m = 0, \end{cases} \quad (4.7)$$

where the normalization comes from the condition  $\langle 0|0\rangle = 1$ .

By solving Eq.(4.6), we can calculate the density of states of fermions. However, it is not easy to obtain solution for arbitrary  $\lambda$ . Then in the following section, we shall solve Eq.(4.6) by perturbative calculation in powers of  $\lambda$ , for  $\lambda = 0$  corresponds to the case of  $\delta$ -function type white noise.

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<sup>1</sup>As shown in Ref[7], right and left states must be distinguished in the present system because of the non-hermiticity of the Hamiltonian. Then only  $\phi_{n=0,m}$  contributes to the norm.

## 5 Solution

In this section we shall obtain wave function of the vacuum state by solving (4.2). We consider the case in which  $\lambda$  is small. The limit  $\lambda \rightarrow 0$  corresponds to the short-range limit which is discussed in Ref.[7].

In order to see the correspondence between the transfer Hamiltonian of the present system (3.5) and that of the short-range limit studied in Ref.[7],

$$H_0 = 2\omega \mathcal{J}^z + 2m_0 \mathcal{J}^x - 2(\mathcal{J}^x)^2, \quad (5.1)$$

we rewrite the Hamiltonian (3.5) as follows,

$$H = H_0 + \frac{1}{\lambda}(a^\dagger + \sqrt{2\lambda}\mathcal{J}^x)(a + \sqrt{2\lambda}\mathcal{J}^x) + \frac{1}{2\lambda}. \quad (5.2)$$

From (5.1), it is easily seen that the second term in Eq.(5.2) should vanish in the  $\lambda \rightarrow 0$ .

The ground state  $|0\rangle$  of the transfer Hamiltonian (3.5) will be obtain by the perturbative calculation from the state  $|0\rangle^{(0)}$ , which is annihilated by the operator  $(a + \sqrt{2\lambda}\mathcal{J}^x)$ ,

$$(a + \sqrt{2\lambda}\mathcal{J}^x)|0\rangle^{(0)} = 0. \quad (5.3)$$

Decomposing the state  $|0\rangle^{(0)}$  into the states  $|n\rangle_0|m\rangle_H$ , namely,  $|0\rangle^{(0)} = \sum \phi_{n,m}^{(0)}|n\rangle_0|m\rangle_H$ , Eq.(5.3) leads to equations of  $\phi_{n,m}^{(0)}$ ,

$$\sqrt{m+1}\phi_{n,m+1}^{(0)} + \sqrt{\frac{\lambda}{2}}n(\phi_{n-1,m}^{(0)} - \phi_{n+1,m}^{(0)}) = 0. \quad (5.4)$$

Equation (5.4) shows that, in the limit  $\lambda \rightarrow 0$ ,  $\phi_{n,m}^{(0)} = 0$  for  $m \geq 1$  and therefore  $\phi_{n,0}^{(0)} = \phi_n$ , where  $\phi_n$  is the wave function for  $\lambda \rightarrow 0$  defined in Ref.[7], i.e.,

$$H_0 \sum_n \phi_n |n\rangle_0 = 0. \quad (5.5)$$

We shall consider higher-order corrections of finite  $\lambda$ . Equation (5.4) can be easily solved as

$$\phi_{n,m}^{(0)} = \frac{1}{\sqrt{m!}}(2\lambda)^{\frac{m}{2}}(n\Delta_n)^m \phi_{n,0}^{(0)}, \quad (5.6)$$

where  $\Delta_n$  is the difference operator with respects to the index  $n$ , i.e.,  $\Delta_n \phi_{n,m}^{(0)} = \frac{\phi_{n+1,m}^{(0)} - \phi_{n-1,m}^{(0)}}{2}$ , and this tells us the order of  $\phi_{n,m}^{(0)}$ ;

$$\phi_{n,m}^{(0)} \sim \lambda^{\frac{m}{2}}. \quad (5.7)$$

Now let us turn to the estimation of the difference between the states  $|0\rangle$  and  $|0\rangle^{(0)}$ .

We decompose the former as

$$|0\rangle = \sum_{n,m} (\phi_{n,m}^{(0)} + \delta\phi_{n,m}) |n\rangle_0 |m\rangle_H, \quad (5.8)$$

where  $\delta\phi_{n,m}$  comes from difference between  $|0\rangle$  and  $|0\rangle^{(0)}$ . Substituting (5.8) into (4.1)

we obtain relation between  $\phi_{n,m}^{(0)}$  and  $\delta\phi_{n,m}$ ,

$$\sum_{n,m} (H_0 \phi_{n,m}^{(0)} + H_\lambda \delta\phi_{n,m}) |n\rangle_0 |m\rangle_H = 0, \quad (5.9)$$

where  $H_\lambda \equiv H - \frac{1}{2\lambda}$ . Noting  $H_\lambda |0\rangle^{(0)} = H_0 |0\rangle^{(0)}$ , we rewrite (5.9) as follows,

$$\begin{aligned} & \sum_{n,m} \left( H_0 \phi_{n,m}^{(0)} + (H_0 + 2(\mathcal{J}^x)^2 + \frac{m}{\lambda}) \delta\phi_{n,m} \right. \\ & \left. + \sqrt{\frac{2}{\lambda}} \mathcal{J}^x (\sqrt{m} \delta\phi_{n,m-1} + \sqrt{m+1} \delta\phi_{n,m+1}) \right) |n\rangle_0 |m\rangle_H = 0. \end{aligned} \quad (5.10)$$

For each  $m$ , the terms should be the same order of  $\lambda$ , i.e.,

$$H_0 \phi_{nm}^{(0)} \sim \frac{1}{\lambda} \delta\phi_{n,m} \sim \frac{1}{\sqrt{\lambda}} (\delta\phi_{n,m-1} + \delta\phi_{n,m+1}). \quad (5.11)$$

From the fact that  $H_0$  does not depend on  $\lambda$  and (5.7), we obtain

$$\delta\phi_{n,m} \sim \lambda^{(\frac{m}{2}+1)}, \quad (5.12)$$

which is of higher order of  $\lambda$  than  $\phi_{n,m}^{(0)}$  for the same  $m$ . From the above result, we can estimate order of the term depending on  $\lambda$  in (5.2). From Eqs.(5.7), (5.12) and the following equation,

$$\begin{aligned} & (a^\dagger + \sqrt{2\lambda} \mathcal{J}^x)(a + \sqrt{2\lambda} \mathcal{J}^x) |0\rangle \\ & = \sum_{n,m} (m \delta\phi_{n,m} + \sqrt{2\lambda} n \Delta_n (\sqrt{m} \delta\phi_{n,m-1} + \sqrt{m+1} \delta\phi_{n,m+1}) \\ & \quad + 2\lambda (n \Delta_n)^2 \delta\phi_{n,m}) |n\rangle_0 |m\rangle_H, \end{aligned} \quad (5.13)$$

we find that the above term is  $\mathcal{O}(\lambda^{\frac{3}{2}})$ . Then it is clear that our estimation is consistent with the expectation that in the limit  $\lambda \rightarrow 0$  the present system becomes that studied in Ref.[7].

In order to calculate the density of states, we have to solve Eq.(4.2). Decomposing it for each powers of  $\lambda$  might be the easiest way to find  $\phi_{n,m}$ .  $\phi_{n,0}^{(0)}$  contains higher-order term of  $\lambda$  than  $\phi_n$ . For  $m = 0$ , Eq.(5.10) reads

$$\sum_n \left( H_0(\phi_{n,0}^{(0)} - \phi_n) + (H_0 + 2(\mathcal{J}^x)^2)\delta\phi_{n,0} + \sqrt{\frac{2}{\lambda}}\mathcal{J}^x\sqrt{m+1}\delta\phi_{n,1} \right) |n\rangle_0 = 0, \quad (5.14)$$

where we use the fact that  $H_0$  annihilates the vacuum for  $\lambda = 0$ , and one can find that  $\phi_{n,0}^{(0)} - \phi_n$  is  $\mathcal{O}(\lambda)$ . Then from (5.6) we can see that  $\phi_{n,m}^{(0)}$  contains a term of  $\mathcal{O}(\lambda^{\frac{m}{2}+1})$  as the next-leading-order term, and therefore we can decompose it as  $\phi_{n,m}^{(0)} = \hat{\phi}_{n,m} + \lambda\check{\phi}_{n,m}$ , and  $\hat{\phi}_{n,m}$  and  $\check{\phi}_{n,m}$  are both  $\mathcal{O}(\lambda^{\frac{m}{2}})$ . We notice the order of  $\lambda$  in Eq.(5.6) and find

$$\hat{\phi}_{n,m} = \frac{1}{\sqrt{m!}}(2\lambda)^{\frac{m}{2}}(n\Delta_n)^m\phi_n. \quad (5.15)$$

We insert the above  $\phi_{n,m}^{(0)}$  into Eq.(5.10) and extract the terms in the order of  $\mathcal{O}(\lambda^{\frac{m}{2}+1})$  and  $\mathcal{O}(\lambda^{\frac{m}{2}})$ ;

$$\begin{aligned} \sum_n \left( \lambda H_0\check{\phi}_{n,m} + (H_0 + 2(\mathcal{J}^x)^2)\delta\phi_{n,m} + \sqrt{\frac{2}{\lambda}}\mathcal{J}^x\sqrt{m+1}\delta\phi_{n,m+1} \right) |n\rangle_0 &= 0, \\ \sum_n \left( H_0\hat{\phi}_{n,m} + \frac{m}{\lambda}\delta\phi_{n,m} + \sqrt{\frac{2}{\lambda}}\mathcal{J}^x\sqrt{m}\delta\phi_{n,m-1} \right) |n\rangle_0 &= 0. \end{aligned} \quad (5.16)$$

By inserting the second equation of (5.16) into the first one, we obtain

$$\begin{aligned} \sum_n H_0(\lambda\check{\phi}_{n,m} + \delta\phi_{n,m})|n\rangle_0 &= \sqrt{\frac{2\lambda}{m+1}} \sum_n \mathcal{J}^x H_0\hat{\phi}_{n,m+1}|n\rangle_0 \\ &= -\frac{2\lambda}{m+1} \sum_n \mathcal{J}^x H_0\mathcal{J}^x\hat{\phi}_{n,m}|n\rangle_0, \end{aligned} \quad (5.17)$$

where we have used Eq.(5.4) in  $\mathcal{O}(\lambda^{\frac{m+1}{2}})$ , namely,

$$\sqrt{m+1} \sum \hat{\phi}_{n,m+1}|n\rangle_0 + \sqrt{2\lambda}\mathcal{J}^x \sum \hat{\phi}_{n,m}|n\rangle_0 = 0.$$

We shall solve Eq.(5.17) from now on.

Let us consider the  $m = 0$  case first, where  $\hat{\phi}_{n,0}$  equals  $\phi_n$ . Appendix C in Ref.[7] gives us the solution  $\phi_n$ . We review it here for completeness. We recall that  $\sum_n \phi_n |n\rangle_0$  is the vacuum state of the transfer Hamiltonian with  $\lambda = 0$ , and thus  $\phi_n$  satisfies the following equation,

$$(2\omega n - 2m_0 n \Delta_n - 2(n \Delta_n)^2) \phi_n = 0. \quad (5.18)$$

If  $\phi(n, m_0)$  is a solution to Eq.(5.18), then so is  $(-1)^n \phi(n, -m_0)$ , and thus general solution can be written as

$$\phi_n = c_1 \phi(n, m_0) + c_2 (-1)^n \phi(n, -m_0), \quad (5.19)$$

with constants  $c_1$  and  $c_2$ . Since  $\phi_n = 1$  and  $\phi_n = (-1)^n$  are solutions of the above equation for  $\omega = m_0 = 0$ ,  $\phi(n, m_0)$  for small  $\omega$  and  $m_0$  is a slowly varying function of  $n$ . This suggests that  $\phi(n, m_0)$  is a “continuum” function of  $n$  and that enables us to replace discrete differences with derivatives. The differential equation for  $\phi(n, m_0)$  is thus obtained as,

$$(2\omega n - 2m_0 n \partial_n - 2(n \partial_n)^2) \phi(n, m_0) = 0. \quad (5.20)$$

In order to solve the above equation, we embed the integer  $n$  into complex number and use the Laplace transformation,

$$\phi(n, m_0) = \int_0^\infty dt e^{-nt} \tilde{\phi}(t). \quad (5.21)$$

By inserting the above Eq.(5.21) into Eq.(5.20), we have

$$\omega \tilde{\phi}(t) - (1 - m_0) t \tilde{\phi}(t) - t^2 \partial_t \tilde{\phi}(t) = 0, \quad (5.22)$$

with the boundary condition  $\tilde{\phi}(t) \rightarrow 0$  in the limit of  $t \rightarrow 0$  or  $t \rightarrow \infty$ . This boundary condition comes from the requirement that  $\phi(n, m_0)$  tends to vanish for large  $n$ . We find the solution

$$\phi(n, m_0) = a_0 \int_0^\infty dt t^{m_0-1} e^{-nt - \frac{\omega}{t}}, \quad (5.23)$$

where  $a_0$  is a constant, and the above solution can be written in terms of the modified Bessel function

$$K_\nu(z) = K_{-\nu}(z) = \frac{1}{2} \left(\frac{z}{2}\right)^\nu \int_0^\infty dt \, e^{-t - \frac{z^2}{4t}} t^{-\nu-1} \quad (5.24)$$

as

$$\phi(n, m_0) = 2a_0 \left(\frac{\omega}{n}\right)^{\frac{m_0}{2}} K_{m_0}(2\sqrt{\omega n}). \quad (5.25)$$

In order to obtain the constants  $c_1 a_0$  and  $c_2 a_0$ , we use the boundary condition [7],

$$\phi_0 \simeq c_1 \phi(1, m_0) + c_2 \phi(1, -m_0) \simeq 1, \quad (5.26)$$

$$\Delta_n \phi_1 \simeq c_1 \partial_n \phi(1, m_0) - c_2 \partial_n \phi(1, -m_0) \simeq 0. \quad (5.27)$$

Please note that the modified Bessel functions have a singularity at  $n = 0$ . Eq.(5.18) is, therefore, ill-defined for  $n = 0$ , but we can expect the “genuine”  $\phi(0, m_0)$  in  $\phi_0$  has almost the same value of  $\phi(1, m_0)$ , since the function  $\phi(n, m_0)$  is assumed to be a slowly varying function. For small  $n$ , we can approximate  $\phi(n, m_0)$  as

$$\begin{aligned} \phi(n, m_0) &= a_0 \int_0^\infty dt \, t^{m_0-1} e^{-nt - \frac{\omega}{t}} \\ &\simeq \frac{a_0}{n^{m_0}} \left( \int_0^{\sqrt{n\omega}} dt \, t^{m_0-1} e^{-\frac{n\omega}{t}} + \int_{\sqrt{n\omega}}^\infty dt \, t^{m_0-1} e^{-t} \right) \\ &\simeq \frac{a_0}{n^{m_0}} (\Gamma(m_0) + (n\omega)^{m_0} \Gamma(-m_0)) \\ &\simeq \frac{a_0}{m_0} (n^{-m_0} - \omega^{m_0}), \end{aligned} \quad (5.28)$$

which coincides with the result obtained in the hard-wall approximation [7]. Using this, Eqs.(5.26) and (5.27) give

$$\frac{c_1 a_0}{m_0} (1 - \omega^{m_0}) + \frac{c_2 a_0}{m_0 \omega^{m_0}} (1 - \omega^{m_0}) = 1, \quad (5.29)$$

$$c_1 a_0 = c_2 a_0. \quad (5.30)$$

Thus the constants are

$$c_1 a_0 = c_2 a_0 = \frac{m_0 \omega^{m_0}}{1 - \omega^{2m_0}}. \quad (5.31)$$

Now let us return to Eq.(5.17). Renumbering the indices of each terms, we have

$$\begin{aligned} (2\omega n - 2m_0 n \Delta_n - 2(n\Delta_n)^2)(\lambda\check{\phi}_{n,m} + \delta\phi_{n,m}) = \\ = -\frac{2\lambda}{m+1}n\Delta_n(2\omega n - 2m_0 n \Delta_n - 2(n\Delta_n)^2)n\Delta_n\hat{\phi}_{n,m}. \end{aligned} \quad (5.32)$$

Since  $\hat{\phi}_{n,m}$  can be decomposed into the form

$$\hat{\phi}_{n,m} = c_1\phi(n, m, m_0) + c_2(-1)^{n+m}\phi(n, m, -m_0), \quad (5.33)$$

we find that  $\lambda\check{\phi}_{n,m} + \delta\phi_{n,m}$  can be also written as

$$\lambda\check{\phi}_{n,m} + \delta\phi_{n,m} = \lambda(c_1\psi(n, m, m_0) + c_2(-1)^{n+m}\psi(n, m, -m_0)). \quad (5.34)$$

As it is for  $\phi_n$ , we can assume that  $\psi(n, m, m_0)$  is a slowly varying function with respect to  $n$ . This again allows us to replace the difference operator  $\Delta_n$  with the differential operator  $\partial_n$ . Eq.(5.32) can be rewritten in the simpler form;

$$\begin{aligned} (2\omega n - 2m_0 n \partial_n - 2(n\partial_n)^2)(\psi(n, m, m_0) + \frac{2}{m+1}(n\partial_n)^2\phi(n, m, m_0)) \\ = -\frac{2}{m+1}2\omega n^2\partial_n\phi(n, m, m_0). \end{aligned} \quad (5.35)$$

We explicitly solve the above equation for the case  $m = 0$ . For  $n \neq 0$ ,

$$\begin{aligned} (2\omega - 2(1 + m_0)\partial_n - 2n(\partial_n)^2)(\psi(n, m_0) + 2(n\partial_n)^2\phi(n, m_0)) = \\ = -4\omega n\partial_n\phi(n, m_0), \end{aligned} \quad (5.36)$$

where  $\psi(n, m_0) = \psi(n, 0, m_0)$ . We again employ Laplace transformation to solve the above equation as before,

$$\psi(n, m_0) + 2(n\partial_n)^2\phi(n, m_0) = \int_0^\infty dt e^{-nt}\tilde{\psi}(t). \quad (5.37)$$

From (5.37) and (5.21), Eq.(5.36) is rewritten as follows;

$$\begin{aligned} (\omega - (1 - m_0)t - t^2\partial_t)\tilde{\psi}(t) &= 2\omega\partial_t(t\tilde{\phi}(t)) \\ &= 2a_0\omega\partial_t(t^{m_0}e^{-\frac{\omega}{t}}) \\ &= 2a_0\omega(m_0t^{m_0-1} + \omega t^{m_0-2})e^{-\frac{\omega}{t}}. \end{aligned} \quad (5.38)$$

In general, we can put  $\tilde{\psi}(t) = C(t)e^{-\frac{\omega}{t}}$ . Inserting this into Eq.(5.38), we find that  $C(t)$  satisfies the following equation,

$$t^2 \partial_t C(t) + (1 - m_0)tC(t) = -2a_0\omega(m_0 t^{m_0-1} + \omega t^{m_0-2}). \quad (5.39)$$

A special solution of this equation is  $C(t) = 2a_0\omega m_0 t^{m_0-2} + a_0\omega^2 t^{m_0-3}$ , and thus we find

$$\begin{aligned} \psi(n, m_0) &= -2(n\partial_n)^2 \phi(n, m_0) + a_0\omega \int_0^\infty dt (2m_0 t^{m_0-2} + \omega t^{m_0-3}) e^{-nt - \frac{\omega}{t}} \\ &= -4a_0(n\partial_n)^2 \left( \left( \frac{\omega}{n} \right)^{\frac{m_0}{2}} K_{m_0}(2\sqrt{n\omega}) \right) \\ &\quad + 2a_0\omega \left( 2m_0 \left( \frac{n}{\omega} \right)^{\frac{1-m_0}{2}} K_{1-m_0}(2\sqrt{n\omega}) + \omega \left( \frac{n}{\omega} \right)^{\frac{2-m_0}{2}} K_{2-m_0}(2\sqrt{n\omega}) \right). \end{aligned} \quad (5.40)$$

Then we find that the general solution to Eq.(5.32) for  $m = 0$  is given by

$$\begin{aligned} \lambda \check{\phi}_{n,0} + \delta \phi_{n,0} &= f \lambda \phi_n - 2\lambda (n\Delta_n)^2 \phi_n \\ &\quad + 2\lambda \omega \left( c_1 a_0 \left( 2m_0 \left( \frac{n}{\omega} \right)^{\frac{1-m_0}{2}} K_{1-m_0}(2\sqrt{n\omega}) + \omega \left( \frac{n}{\omega} \right)^{\frac{2-m_0}{2}} K_{2-m_0}(2\sqrt{n\omega}) \right) \right. \\ &\quad \left. + c_2 a_0 (-1)^n \left( -2m_0 \left( \frac{n}{\omega} \right)^{\frac{1+m_0}{2}} K_{1+m_0}(2\sqrt{n\omega}) + \omega \left( \frac{n}{\omega} \right)^{\frac{2+m_0}{2}} K_{2+m_0}(2\sqrt{n\omega}) \right) \right), \end{aligned} \quad (5.41)$$

where  $f$  is a constant which is to be determined by the boundary condition Eq.(4.7). For regular  $\phi_n$ 's,  $(n\Delta_n)^m \phi_n = 0$  for  $n = 0$ , since this term is proportional to  $n$ . This gives  $\hat{\phi}_{0,m} = 0$  for  $m > 0$ , which can be verified by practical calculation. Using  $\phi_0 = 1$ , the condition Eq.(4.7) requires

$$\lambda \check{\phi}_{0,m} + \delta \phi_{0,m} = 0, \quad (5.42)$$

for arbitrary  $m$ . We insert (5.41) into (5.42) for the case  $m = 0$  in order to determine the constant  $f$ . For small  $n$ , the modified Bessel function is approximated as

$$2 \left( \frac{n}{\omega} \right)^{\frac{1 \pm m_0}{2}} K_{1 \pm m_0}(2\sqrt{n\omega}) \simeq \omega^{\mp m_0 - 1}, \quad (5.43)$$

$$2 \left( \frac{n}{\omega} \right)^{\frac{2 \pm m_0}{2}} K_{2 \pm m_0}(2\sqrt{n\omega}) \simeq \omega^{\mp m_0 - 2}. \quad (5.44)$$



Then Eq.(5.42) for  $m = 0$  is

$$\begin{aligned}
& \lambda \check{\phi}_{0,0} + \delta \phi_{0,0} \\
&= f\lambda + \lambda(c_1 a_0 \omega(2m_0 \omega^{m_0-1} + \omega \omega^{m_0-2}) + c_2 a_0 \omega(-2m_0 \omega^{-m_0-1} + \omega \omega^{-m_0-2})) \\
&= f\lambda - \lambda \frac{m_0 \omega^{m_0}}{1 - \omega^{2m_0}} (2m_0 \frac{1 - \omega^{2m_0}}{\omega^{m_0}} - \frac{1 + \omega^{2m_0}}{\omega^{m_0}}) \\
&= f\lambda - 2\lambda m_0^2 + \lambda m_0 \frac{1 + \omega^{2m_0}}{1 - \omega^{2m_0}} = 0.
\end{aligned} \tag{5.45}$$

Thus

$$\phi_{n,0} = \phi_n + \lambda \psi_n, \tag{5.46}$$

where

$$\begin{aligned}
\psi_n &= (2m_0^2 - m_0 \frac{1 + \omega^{2m_0}}{1 - \omega^{2m_0}}) \phi_n - 2(n\Delta_n)^2 \phi_n \\
&+ 2\omega \frac{m_0 \omega^{m_0}}{1 - \omega^{2m_0}} \left( 2m_0 \left(\frac{n}{\omega}\right)^{\frac{1-m_0}{2}} K_{1-m_0}(2\sqrt{n\omega}) + \omega \left(\frac{n}{\omega}\right)^{\frac{2-m_0}{2}} K_{2-m_0}(2\sqrt{n\omega}) \right. \\
&\left. + (-1)^n (-2m_0 \left(\frac{n}{\omega}\right)^{\frac{1+m_0}{2}} K_{1+m_0}(2\sqrt{n\omega}) + \omega \left(\frac{n}{\omega}\right)^{\frac{2+m_0}{2}} K_{2+m_0}(2\sqrt{n\omega})) \right).
\end{aligned} \tag{5.47}$$

The other components  $\phi_{n,m}$  for  $m \geq 1$  can be calculated in a similar way to the above. Calculations are summarized in Appendix A.

## 6 The density of states

Now we can calculate the density of states. The density of states is obtained from  $\bar{G}_{\alpha\beta}$  as

$$\rho(\epsilon) = \frac{1}{\pi} \lim_{\omega \rightarrow 0} \text{Im} \bar{\mathcal{G}}(0; \epsilon + i\omega), \tag{6.1}$$

where the averaged Green's function  $\bar{\mathcal{G}}(x; i\omega)$  is given by,

$$\bar{\mathcal{G}}(x; i\omega) = i^x \sum_{\alpha} (-1)^{\alpha x} \bar{G}_{\alpha\alpha}(x; i\omega), \tag{6.2}$$

with  $\alpha = 0(1)$  for  $R(L)$ . As shown in Ref.[7],

$$\bar{\mathcal{G}}(0; i\omega) = i \sum_{n,m} (-1)^n |\phi_{n,m}|^2, \tag{6.3}$$

and we shall calculate  $\rho(\epsilon)$  by performing analytic continuation  $i\omega \rightarrow \epsilon + i\omega$ .

The low-order corrections of  $\lambda$  in the averaged Green function is given as

$$\begin{aligned}
\bar{\mathcal{G}}(0; i\omega) &= i \sum_{n=0}^{\infty} (-1)^n \phi_{n,0}^2 + i \sum_{n=0}^{\infty} (-1)^n \phi_{n,1}^2 \cdots \\
&= i \sum_n (-1)^n \phi_n^2 + \\
&\quad + 2i \sum_n (-1)^n \phi_n (\lambda \check{\phi}_{n,0} + \delta \phi_{n,0}) + i \sum_n (-1)^n \hat{\phi}_{n,1}^2 \\
&\quad + \mathcal{O}(\lambda^2),
\end{aligned} \tag{6.4}$$

where each row corresponds to the order  $\mathcal{O}(1)$ ,  $\mathcal{O}(\lambda)$ , and higher. Inserting the  $\hat{\phi}_{n,m}$  and  $\lambda \check{\phi}_{n,m} + \delta \phi_{n,m}$ , and using the formula [11]

$$\begin{aligned}
&\int_0^{\infty} \mathrm{d}n \left( 2 \left( \frac{n}{\omega} \right)^{\frac{\mu}{2}} K_{\mu}(2\sqrt{n\omega}) \cdot 2 \left( \frac{n}{\omega} \right)^{\frac{\nu}{2}} K_{\nu}(2\sqrt{n\omega}) \right) \\
&= \int_0^{\infty} \mathrm{d}z \frac{z}{2\omega} \cdot 2 \left( \frac{z}{2\omega} \right)^{\mu} K_{\mu}(z) \cdot 2 \left( \frac{z}{2\omega} \right)^{\nu} K_{\nu}(z) \\
&= 4 \left( \frac{1}{2\omega} \right)^{\mu+\nu+1} \cdot \frac{2^{\mu+\nu-1}}{1+\mu+\nu} \Gamma(1+\mu) \Gamma(1+\nu),
\end{aligned} \tag{6.5}$$

we have

$$\begin{aligned}
\sum_n (-1)^n \phi_n^2 &\simeq 2c_1 c_2 \int_0^{\infty} \mathrm{d}n \phi(n, m_0) \phi(n, -m_0) \\
&\simeq \frac{2m_0^2 \omega^{2m_0}}{\omega(1 - \omega^{2m_0})^2}
\end{aligned} \tag{6.6}$$

$$\sum_n (-1)^n \phi_n (\lambda \check{\phi}_{n,0} + \delta \phi_{n,0}) \simeq -\frac{\lambda m_0^2 \omega^{2m_0}}{(1 - \omega^{2m_0})^2} \left( \frac{2m_0(1 + \omega^{2m_0})}{\omega(1 - \omega^{2m_0})} - \frac{2}{3\omega} \right) \tag{6.7}$$

$$\sum_n (-1)^n \hat{\phi}_{n,1}^2 \simeq \frac{2\lambda m_0^2 \omega^{2m_0}}{(1 - \omega^{2m_0})^2} \left( \frac{2m_0^2}{\omega} - \frac{2}{3\omega} \right). \tag{6.8}$$

Higher-order terms can be calculated in a similar way. We are interested in the averaged Green's function in the limit  $m_0 \rightarrow 0$ , since the delocalization transition occurs there. Collecting the higher-order terms,

$$\bar{\mathcal{G}}(0; i\omega) = \frac{i}{2\omega(\ln \omega)^2}$$

$$\begin{aligned}
& +\lambda\left(\frac{i}{\omega(\ln \omega)^3}\right) \\
& +\lambda^2\left(\frac{13i}{30\omega(\ln \omega)^2}+\frac{i}{3\omega(\ln \omega)^3}+\frac{i}{2\omega(\ln \omega)^4}\right) \\
& -\lambda^3\frac{1}{210}\frac{i}{\omega(\ln \omega)^2} \\
& +\mathcal{O}(\lambda^4).
\end{aligned} \tag{6.9}$$

Each term in the averaged Green's function has the form  $\frac{i}{\omega(\ln \omega)^\alpha}$ . We perform analytic continuation in order to obtain the density of states  $\rho(\epsilon)$  such as

$$\text{Im}\frac{i}{\omega(\ln \omega)^\alpha} \rightarrow -\text{Im}\frac{1}{\epsilon(\ln \epsilon - i\frac{\pi}{2})^\alpha} \simeq (-1)^\alpha \frac{\alpha\pi}{2\epsilon|\ln \epsilon|^{\alpha+1}}, \tag{6.10}$$

for small  $\epsilon$ . Recovering the constant  $g$ , the density of states is obtained as follows,

$$\begin{aligned}
\rho(\epsilon) = & \frac{1}{2\frac{\epsilon}{2g}(\ln \frac{\epsilon}{2g})^3} \\
& -2g\lambda\frac{3}{2\frac{\epsilon}{2g}(\ln \frac{\epsilon}{2g})^4} \\
& +4g^2\lambda^2\left(\frac{13}{30\frac{\epsilon}{2g}(\ln \frac{\epsilon}{2g})^3}-\frac{1}{2\frac{\epsilon}{2g}(\ln \frac{\epsilon}{2g})^4}+\frac{1}{\frac{\epsilon}{2g}(\ln \frac{\epsilon}{2g})^5}\right) \\
& -8g^3\lambda^3\frac{1}{210\frac{\epsilon}{2g}(\ln \frac{\epsilon}{2g})^3} \\
& +\mathcal{O}((g\lambda)^4).
\end{aligned} \tag{6.11}$$

The above expression (6.11) is the main result of this paper.

## 7 Discussion

In the previous sections, we have obtained density of states of fermions interacting with quenched disorder of nonlocal correlation. The result is given by (6.11) and Figs.1 and 2. It is obvious that unit of the energy  $\epsilon$  is  $g$ , the magnitude of disorder fluctuation, and  $\rho(\epsilon)$  scales with  $(g\lambda)$ . In the telegraph process of random disorder[6],  $\frac{g}{\lambda}$  and  $\lambda$  correspond to the mean height and width of fluctuating mass  $m(x)$ , respectively.

Recently we studied the eigenvalue problem

$$h\psi = E\psi \tag{7.1}$$

by numerical calculation for telegraphic  $m(x)$ , and found that all the eigenvalues are determined by the combination  $(g\lambda)$  in unit  $g$ [12].

In Fig.1, we show  $\rho(\epsilon)$  for various  $\lambda$  with a fixed  $g$ . It is easily seen that for  $\epsilon$  smaller than  $\epsilon_0 \sim ge^{-\frac{1}{g\lambda}}$ ,  $\rho(\epsilon)$  is a increasing function of  $\lambda$  whereas for  $\epsilon > \epsilon_0$  it is a decreasing function. That is, states near the band center are enhanced by the nonlocal correlation of disorder, which partially suppresses random fluctuation of disorder. From the above result, we expect that the number of extended states near the band center is increased by the existence of such kind of disorder correlation.

The above expectation can be verified by solving (7.1). In order to study the field equation (7.1) in the multi-solitonic background of  $m(x)$ , transfer-matrix formalism is very useful[12]. Using this method, we have obtained solutions to (7.1) in various configurations of  $m(x)$ . We think that from these explicit form of solutions we can understand characteristic properties of the present random system and their origins, e.g., power-law decay of Green's functions, complex multi-fractal scaling[13], etc.

Results in [12] are summarized as follows;

1. For quasi-periodic  $m(x)$  almost all states are extended and quasi-periodic. Peaks of the wave functions appear quasi-periodically. This is a reminiscence of Bloch's theorem although we are *not* imposing the periodic boundary condition on the wave function. This result remains correct even for  $m_0 = \langle m(x) \rangle \neq 0$ .
2. For large random fluctuation of disorder close to the white noise, almost all states tend to "localized" even for small or vanishing  $m_0$ . They are divided into two categories; the first one simply has one peak or peaks close each other, whereas the second one has more than two peaks which are spatially separated comparatively large distance. The first one is obviously genuine localized state.

3. On suppressing randomness of disorder, the number of states of the second category increases.

We think that the above results support our conclusion from the density of states, i.e., the second category is main part of extended states and the number of extended states is increased by nonlocal correlation of disorder. We also expect that on increasing correlation of disorder, quasi-periodic states in quasi-periodic  $m(x)$  come to contribute to physical quantities as extended state.

Sometimes properties of the present disordered system are explained by using the single zero-energy wave function  $\psi_{\pm} \propto e^{\pm \int^x dx' m(x')}$  [14]. However as the behavior of the density of states shows, a large number of low-energy states are contributing to physical quantities. Then it is important to see if these “typical” low-energy modes have almost similar behaviors to those of  $\psi_{\pm}$ . Details will be reported in a future publication[12]. We also calculate Green’s functions in the present system[10]. All these results are quite useful for understanding localization and delocalization transition.

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## Appendix

### A Higher-order terms

Higher-order corrections to the averaged Green's function are calculated similarly to the correction of  $\mathcal{O}(\lambda)$ . It is easy to see that the limit,  $m_0 \rightarrow 0$ , and the integration, Eq.(6.5), are commutable with each other, and thus we take the limit  $m_0 \rightarrow 0$  first.

From the definition of  $\hat{\phi}_{n,m}$ , Eq.(5.15), and taking that limit, we have

$$\hat{\phi}_{n,m} = -\frac{1}{2a_0 \ln \omega} \phi(n, m, 0) (1 + (-1)^{n+m}) \quad (m = 1, 2, 3, \dots), \quad (\text{A.1})$$

with

$$\begin{aligned} \phi(n, 1, 0) &= -a_0 \sqrt{2\lambda} \omega \cdot 2\left(\frac{n}{\omega}\right)^{\frac{1}{2}} K_1(2\sqrt{n\omega}), \\ \phi(n, 2, 0) &= \sqrt{2} a_0 \lambda \left( -\omega \cdot 2\left(\frac{n}{\omega}\right)^{\frac{1}{2}} K_1(2\sqrt{n\omega}) + \omega^2 \cdot 2\left(\frac{n}{\omega}\right)^{\frac{1}{2}} K_1(2\sqrt{n\omega}) \right), \\ \phi(n, 3, 0) &= -\frac{2}{\sqrt{3}} a_0 \lambda^{\frac{3}{2}} \left( \omega \cdot 2\left(\frac{n}{\omega}\right)^{\frac{1}{2}} K_1(2\sqrt{n\omega}) - 3\omega^2 \cdot 2\left(\frac{n}{\omega}\right) K_2(2\sqrt{n\omega}) \right. \\ &\quad \left. + \omega^3 \cdot 2\left(\frac{n}{\omega}\right)^{\frac{3}{2}} K_3(2\sqrt{n\omega}) \right). \end{aligned}$$

Solving Eq.(5.32), we have the higher-order terms in  $\phi_{n,m}$

$$\lambda \check{\phi}_{n,m} + \delta \phi_{n,m} = -\frac{1}{2a_0 \ln \omega} \psi(n, m, 0) (1 + (-1)^{n+m}) \quad (m = 1, 2, 3, \dots), \quad (\text{A.2})$$

with

$$\begin{aligned} \psi(n, 1, 0) &= -\sqrt{2} a_0 \lambda^{\frac{3}{2}} \left( \omega \cdot 2\left(\frac{n}{\omega}\right)^{\frac{1}{2}} K_1(2\sqrt{n\omega}) - \frac{5}{2} \omega^2 \cdot 2\left(\frac{n}{\omega}\right) K_2(2\sqrt{n\omega}) \right. \\ &\quad \left. + \frac{2}{3} \omega^3 \cdot 2\left(\frac{n}{\omega}\right)^{\frac{3}{2}} K_3(2\sqrt{n\omega}) \right) \\ \psi(n, 2, 0) &= \frac{2\sqrt{2}}{3} a_0 \left( \omega \cdot 2\left(\frac{n}{\omega}\right)^{\frac{1}{2}} K_1(2\sqrt{n\omega}) - \frac{13}{2} \omega^2 \cdot 2\left(\frac{n}{\omega}\right) K_2(2\sqrt{n\omega}) \right. \\ &\quad \left. + 5\omega^3 \cdot 2\left(\frac{n}{\omega}\right)^{\frac{3}{2}} K_3(2\sqrt{n\omega}) - \frac{3}{4} \omega^4 \cdot 2\left(\frac{n}{\omega}\right)^2 K_4(2\sqrt{n\omega}) \right). \end{aligned}$$

From the above results, we have the higher corrections to the averaged Green's function; Terms in  $\mathcal{O}(\lambda)$  are

$$\begin{aligned}\sum_n \hat{\phi}_{n,1}^2 (-1)^n &= -\frac{\lambda}{3} \frac{1}{\omega (\ln \omega)^2}, \\ 2 \sum_n \hat{\phi}_{n,0} (\lambda \check{\phi}_{n,0} + \delta \phi_{n,0}) (-1)^n &= \lambda \left( \frac{1}{3} \frac{1}{\omega (\ln \omega)^2} + \frac{1}{\omega (\ln \omega)^3} \right),\end{aligned}$$

terms in  $\mathcal{O}(\lambda^2)$  are

$$\begin{aligned}\sum_n \hat{\phi}_{n,2}^2 (-1)^n &= \lambda^2 \frac{2}{15} \frac{1}{\omega (\ln \omega)^2}, \\ 2 \sum_n \hat{\phi}_{n,1} (\lambda \check{\phi}_{n,1} + \delta \phi_{n,1}) (-1)^n &= \lambda^2 \left( \frac{7}{30} \frac{1}{\omega (\ln \omega)^2} \right), \\ \sum_n (\lambda \check{\phi}_{n,0} + \delta \phi_{n,0})^2 (-1)^n &= \lambda^2 \left( \frac{1}{15} \frac{1}{\omega (\ln \omega)^2} + \frac{1}{3} \frac{1}{\omega (\ln \omega)^3} + \frac{1}{2} \frac{1}{\omega (\ln \omega)^4} \right),\end{aligned}$$

and terms in  $\mathcal{O}(\lambda^3)$  are

$$\begin{aligned}\sum_n \hat{\phi}_{n,3}^2 (-1)^n &= -\lambda^3 \frac{16}{315} \frac{1}{\omega (\ln \omega)^2}, \\ 2 \sum_n \hat{\phi}_{n,2} (\lambda \check{\phi}_{n,2} + \delta \phi_{n,2}) (-1)^n &= \lambda^3 \left( \frac{31}{315} \frac{1}{\omega (\ln \omega)^2} \right), \\ \sum_n (\lambda \check{\phi}_{n,1} + \delta \phi_{n,1})^2 (-1)^n &= -\lambda^3 \left( \frac{11}{210} \frac{1}{\omega (\ln \omega)^2} \right).\end{aligned}$$

Combining these terms, we have the averaged Green function Eq.(6.9).

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## Figure captions

Fig. 1 : The density of states, as a function of the energy, whose  $\lambda$  is varied from 0.00 to 0.80 at  $g = 0.5$ . The axis of abscissa stands for the energy ( $\epsilon$ ) and the one of column for the density of states ( $\rho(\epsilon)$ ).

Fig. 2 : The density of states whose  $g$  is varied from 0.25 to 2.0 at  $\lambda = 0.1$ .



